

## **Outer boundary conditions in General Relativity**

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Outer boundary conditions in GR - p.1/2

## Outline



- Introduction
- Formulation
- Constraint-preserving boundary conditions
- Determinant condition
- Numerical results
- A related toy model problem in ED

## Introduction



Solve Einstein's equations in a domain with timelike boundaries.



#### Boundary conditions should

(i) be compatible with the constraints (constraint-preserving)(ii) be physically reasonable (e.g. minimize reflections)(iii) yield a well posed initial-boundary value formulation

## Introduction



- A well posed initial-boundary value formulation was given by Friedrich & Nagy, 1999 in terms of a tetrad-based Einstein-Bianchi formulation.
- Numerical implementation for related formulation is underway (Reula, Bardeen, Buchman, S,...?)
- Less is known for metric-based formulations (although recent progress by Cornell-Caltech group and S & Tiglio).
- Relevant for: Outer/interface boundary conditions; constraint projection, elliptic gauge conditions,...

## Formulation



Evolution equations can be cast into first order quasilinear form: Frittelli & Reula, Anderson & York, Hern, KST,..., S & Tiglio

$$\pounds_n \alpha = -\alpha K,$$

$$\begin{aligned} \pounds_{n}g_{ij} &= -2K_{ij}, \\ \xi_{n}K_{ij} &= \frac{1}{2}g^{ab}\left(-\partial_{a}d_{bij} + 2\partial_{(i}d_{|ab|j)} - \partial_{(i}d_{j)ab} - 2\partial_{(i}A_{j)}\right) + \gamma g_{ij}H + \mathsf{I.c.} \\ \xi_{n}d_{kij} &= -2\partial_{k}K_{ij} + \eta g_{k(i}M_{j)} + \gamma g_{ij}M_{k} + \mathsf{I.o.} \end{aligned}$$

 $\pounds_n A_i = -KA_i - g^{ab}\partial_i K_{ab} + \xi M_i + \text{l.o.}$ 

with some parameters  $\gamma$ ,  $\eta$ ,  $\chi \xi$ . Constraints: H = 0,  $M_j = 0$  (Hamiltonian and momentum),  $d_{kij} = \partial_k g_{ij}$ ,  $A_i = \partial_i \alpha / \alpha$ .

## Formulation



Main evolution system has the form

$$\partial_t u = P^i(u)\partial_i u + F(u),$$

where  $u = (\alpha, g_{ij}, K_{ij}, d_{kij}, A_k)$ .

The constraint variables  $v = (H, M_j, d_{kij} - \partial_k g_{ij}, A_i - \partial_i \alpha / \alpha, ...)$ satisfy the constraint propagation system

$$\partial_t v = Q^i(u)\partial_i v + B[u]v,$$

Provided that the parameters  $\gamma$ ,  $\eta$ ,  $\chi \xi$  satisfy suitable inequalities, these two systems can be brought into *strongly hyperbolic form*. So in the absence of boundaries we have a well posed formulation.

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Provided that the parameters  $\gamma$ ,  $\eta$ ,  $\chi \xi$  satisfy suitable inequalities, these two systems can be brought into *strongly hyperbolic form*. So in the absence of boundaries we have a well posed formulation.  $\partial_t u = P^i(u)\partial_i u + F(u)$  is called *strongly hyperbolic* if there exists K > 0 and a symmetric matrix-valued function H(u, n) which is smooth in u and n such that  $K^{-1} \leq H(u, n) \leq K$  and  $H(u, n)P^i(u)n_i$  is symmetric for all  $n \in S^2$  and all u.



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- Require this system to be symmetric hyperbolic, i.e. the symmetrizer H(u) = H(u, n) is independent of n.
- Specify maximal dissipative boundary conditions:

$$E(t) \equiv \int_{\Omega} v^T H v \, d^3 x, \qquad \frac{d}{dt} E(t) \leq \int_{\partial \Omega} v^T H Q(n) v \, dS + \frac{1}{\tau} E(t).$$

 $v^T HQ(n)v = v_{in}^T \Lambda_+ v_{in} - v_{out}^T \Lambda_- v_{out}$ . Set  $v_{in} = 0$  (3 b.c.). In this case we have an energy estimate  $E(t) \le e^{t/\tau} E(0)$ . In particular, this implies that v(t) = 0 if v(0) = 0.



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- Boundary matrix H(u,n)P<sup>i</sup>(u)n<sub>i</sub> has six positive eigenvalues; for high-frequency plane waves propagating towards the boundary: three constraint-violating modes; fields u<sup>(cons)</sup><sub>in</sub> two physical modes; fields u<sup>(phys)</sup><sub>in</sub> one gauge mode; fields u<sup>(gauge)</sup><sub>in</sub>



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- Notice:  $u_{in}^{(cons)} \neq v_{in}!$  Rather, the three conditions  $v_{in} = 0$  yield a differential boundary condition for  $u_{in}^{(cons)}$  at the boundary:  $\partial_t u_{in}^{(cons)} = \dots$



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- Set  $u_{in}^{(gauge)} = 0$ .



A different way of specifying boundary data is through the Weyl scalars  $\Psi_0$  and  $\Psi_4$ , constructed from an adapted NP tetrad at the boundary:

 $\Psi_0 = c\Psi_4^* + h.$ 

where |c| < 1.

Notice:

For linear fluctuations about a Schwarzschild black holes and spherically symmetric outer boundary,  $\Psi_0$  and  $\Psi_4$  are gauge-invariant quantities.

## **Determinant condition**



Consider linear hyperbolic system with constant coefficients (high-frequency limit),

 $\partial_t u = \mathcal{A}u, t > 0, x > 0,$ 

where  $Au \equiv A^x \partial_x u + A^y \partial_y u + A^z \partial_z u$  with differential boundary conditions

$$M(\partial_x, \partial_y, \partial_z)u = h(t, y, z).$$

Look for solutions of the form  $u(t, x, y, z) = e^{st+i(w_y y+w_z z)} f(x)$ , where Re(s) > 0,  $w_y$ ,  $w_z$  real.

Test: If h = 0 there should be no such solutions. Otherwise the system is ill posed: Because if there is such a solution for some s, Re(s) > 0, then there is also a solution  $u_{\alpha}$  for  $\alpha s$ ,  $\alpha > 0$  and for each fixed t

$$|u_{\alpha}(t, x, y, z)| / |u_{\alpha}(0, x, y, z)| = e^{\alpha Re(s)t} \to \infty$$

(i.e. the operator s - A is not invertible for all Re(s) > 0.)

## **Determinant condition**



Introducing the ansatz  $u(t, x, y, z) = e^{st+i(w_y y+w_z z)} f(x)$  into the evolution and boundary equations gives

 $sf = A^x \partial_x f + i(A^y w_y + A^z w_z) f, \qquad L(s, iw_y, iw_z) f = 0.$ 

Solution has the form  $f(x) = Pe^{M_-x}\sigma_-$ ,  $Re(M_-) < 0$  with  $LP\sigma_- = 0$ . Therefore, one has to verify the determinant condition

 $\det(LP)(s, w_y, w_z) \neq 0, \qquad Re(s) > 0.$ 

One can rule out "candidate" constraint-preserving boundary conditions (Calabrese, OS, J. Math. Phys. 44, 3888 (2003)). Such ill posed solutions can be constraint-violating or gauge modes!



- 3D numerical finite-difference code (Lehner, Nielsen, Tiglio)
- Domain is a cubic box  $[-1,1]^3$ .
- Third order Runge-Kutta time-discretization.
- Second-order accurate finite differencing for spatial operators.
- Some artificial dissipation.















#### Comparison with non-constraint-preserving boundary conditions







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Fat Maxwell  $(A_i \leftrightarrow g_{ij}, E_j \leftrightarrow K_{ij}, W_{ij} \leftrightarrow d_{kij})$ :

$$\partial_t A_i = E_i + \nabla_i \phi,$$
  

$$\partial_t E_j = \nabla^i (W_{ij} - W_{ji}) + \alpha \delta^{ij} C_{kij},$$
  

$$\partial_t W_{ij} = \nabla_i E_j + \frac{\beta}{2} \delta_{ij} \rho + \nabla_i \nabla_j \phi,$$

with the constraints  $\rho \equiv \nabla^k E_k = 0$ ,  $C_{kij} = \nabla_k W_{ij} - \nabla_i W_{kj} = 0$ . Strongly hyperbolic if  $\alpha\beta > 0$  (Cauchy problem well posed in  $L^2$ ). If boundaries are present, impose the boundary conditions

 $\nabla^k E_k = 0$  preserves the constraints

 $\mathbf{E}_{||} = (W_{n||} - W_{||n}) + h_{||}$ 

controls normal component of Poynting vecto



Choose the gauge condition  $\phi = 0$  (temporal gauge  $\leftrightarrow$  fixed shift).

• Well posed in  $L^2$  ( $u = (A_i, \overline{E_j, W_{ij}})$ )?

$$\|u(t,.)\|_{L^{2}(\Omega)} \leq ae^{bt} \left[ \|u(0,.)\|_{L^{2}(\Omega)} + \int_{0}^{t} \|h(s)\|_{L^{2}(\partial\Omega)} ds \right].$$



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- The system passes the determinant condition for all  $\alpha\beta > 0$
- However, consider solutions of the type

$$A_i = t \nabla_i f, \qquad E_j = \nabla_i f, \qquad W_{ij} = t \nabla_i \nabla_j f,$$

where f is a smooth, time-independent, harmonic function. Evolution and constraints equations are satisfied. Initial and boundary data only depend on first derivatives of f whereas the solution depends on second derivatives of f.



- This is due to a bad gauge choice at the boundary! (physically one has an electrostatic solution with nontrivial electric charge density at the boundary)
- This motivates the following gauge condition:

$$\Delta \phi = -\nabla^k E_k$$
, on boundary:  $\partial_n \phi = -E_n$ .

Using this gauge condition, one can show that the problem is well posed in a Hilbert space that controls the  $L^2$  norm of the fields *and* the constraint variables (Reula, S, gr-qc/0409027) Solution flux in this space is given by a semigroup.

Current work with G. Nagy for generalization to Einstein (maximal slicing and minimal strain).

## Conclusions



At the end of the day ???

Matching to a characteristic code (Bishop, Winicour, d'Inverno, ...)



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Conformal field equations (Friedrich et al.)

